The entrance channel effect on the synthesis of a superheavy element ²⁹⁶119

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Abstract: In this study, we investigated the entrance channel effect on the evaporation residue cross section of a superheavy element ²⁹⁶119. Using 29 projectile-target combinations, we investigated the role of the entrance channel on the 3*n* and 4*n* evaporation channels in hot combinations. This effect can be evaluated based on the entrance channel asymmetry and Q value of complete fusion. We calculated the variation of the maximum evaporation residue cross sections (σ_{3n}^{\max} and σ_{4n}^{\max}) with |Q| for the reactions ⁴⁹⁻⁴⁷Ti+²⁴⁷⁻²⁴⁹Bk, ⁶⁰⁻⁵⁷Fe+²³⁶⁻²³⁹Np, ⁴⁴⁻⁴²Ca+²⁵²⁻²⁵⁴Es, and ^{55,54,52}Mn+^{241,242,244}Pu. With an increase in |Q|, σ_{3n}^{\max} and σ_{4n}^{\max} increase. In addition, we studied the role of asymmetry and mean fissility parameters in the synthesis of the superheavy element. The obtained results in this study can be utilized in future studies.

Keywords: evaporation residue cross section, entrance channel, superheavy element

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I. INTRODUCTION

The study on the formation of superheavy elements via heavy ion fusion reactions is an interesting topic in nuclear physics. Heavy ion fusion reactions can be adopted to study the different properties of the compound nucleus [1-5]. Using projectile and target nuclei, the compound nucleus can be formed. To obtain a heavy nucleus, a heavy projectile is required to be fused with a heavy target. There are two experimental methods for synthesizing superheavy elements. The cold fusion method is used to synthesize superheavy elements with Z = 107-112 [6, 7]. Another method is the hot fusion reaction method, which is used to produce elements with Z = 113-118 [8-12]. It should be noted that an element with Z = 113 was also produced via the cold fusion reaction [13, 14].

For the synthesis of a superheavy element with Z = 119, Zagrebaev and Greiner experimentally reported the upper cross sectional 50 fb in the reaction ⁵⁰Ti+²⁴⁹Bk, and no indication for the discovery of an element with Z = 119 was observed. Recently, Khuyagbaatar et al. [15] investigated the production of superheavy elements with Z = 119 and 120 in the reactions ⁵⁰Ti+²⁴⁹Bk and ⁵⁰Ti+²⁴⁹Cf. They discussed the nonobservation of the elements (with Z = 119 and 120) using the concept of fusion evaporation reactions, addressing different theoretical calculations on the fission barrier heights of superheavy elements. Different theoretical works exist with different models such as the multidimensional stochastic [16, 17], dinuclear system [18, 19], and fusion-by-diffu-

sion models [20, 21]. In Ref. [22], within the dinuclear system model, the production of hassium isotopes is studied in the reactions $^{22}Ne + ^{249}Cf$, $^{25,26}Mg + ^{248}24Cm$, $^{30}Si + ^{244}Pu$, $^{34,36}S + ^{238}U$, $^{40}Ar + ^{232}Th$, and $^{48}Ca + ^{226}Ra$.

It is clear that synthesizing superheavy nuclei is very difficult experimentally because the evaporation residue cross section of these reactions is negligible. Consequently, the study on the effective parameters of evaporation residue cross section is crucial. These parameters are different, and they include incident energy, combination of the projectile and target, deformation of the colliding nuclei, etc. Recently, we calculated the evaporation residue cross section for superheavy elements with Z = 120, considering the projectile and target deformations [23].

In heavy ions collision, the quasifission processes occur on short time scales at approximately 10^{-20} s [24-26], while fusion-fission usually occurs on longer time scales, from approximately 10^{-20} to 10^{-16} s. There is a close relationship between the production of superheavy elements and entrance channels. In other words, the effects of entrance channel and the competition between the quasifission processes and complete fusion can facilitate the elucidation of heavy ion collisions. This subject has been studied in several publications [27-30]. In addition, several studies exist on the influence of the neutron numbers of projectile and target on the evaporation residue cross sections in hot fusion reactions [31-34].

In this study, using the Coulomb and proximity potential, we investigated the role of the entrance channel of

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colliding nuclei on the evaporation residue cross section of a superheavy element ²⁹⁶119 formed in heavy ion fusion reactions. We also investigated the relationship between the mean fission parameter, quasi-fission, and evaporation residue cross section. This paper is organized as follows. In Sec. II, the theoretical method is provided. Then, the obtained results are presented in Sec. III. Finally, the summary and conclusion of our results are provided in Sec. IV.

II. THEORETICAL METHOD

A. The potential

The potential between the target and the projectile can be calculated based on the Coulomb and proximity potential model as [35]

$$V(r) = V_c + V_P + \frac{\hbar^2 l(l+1)}{2\mu r^2},$$
(1)

where, l is the angular momentum, r represents the distance between the centers of the projectile and target, and μ is the reduced mass. The Coulomb potential is given as

$$V_c = \frac{Z_P Z_T e^2}{r},\tag{2}$$

where, Z_P and Z_T are the atomic numbers of the projectile and target, respectively.

 V_P is the proximity potential given by [36]

$$V_P = 4\pi\gamma b \left(\frac{C_1 C_2}{C_1 + C_2}\right) \Phi(z/b),\tag{3}$$

where z is the short separation distance and the nuclear surface thickness is given as $b \approx 1$ fm. In addition, the surface energy constant is expressed as

$$\gamma = 0.9517 \left[1 - 1.7826 \left(\frac{N-Z}{A} \right)^2 \right] \text{MeV} \text{fm}^{-2},$$
 (4)

where, N, Z, and A represent the neutron, proton, and atomic numbers of the parent nucleus, respectively. C_i can be calculated as

$$C_i = R_i - \frac{b^2}{R_i^2},\tag{5}$$

with

$$R_i = 1.24A_i^{1/3} - 0.76 + 0.8A_i^{-1/3}.$$
 (6)

The universal function $\Phi(z/b)$ can be calculated as [36]

$$\Phi(z/b) = \begin{cases} -1.7817 + 0.9270z + 0.0143z^2 - 0.09z^3 & z \le 0, \\ -1.7817 + 0.9270z + 0.01696z^2 - 0.05148z^3 & 0 < z < 1.9475, \\ -4.41\exp(-z/0.7176) & z \ge 1.9475. \end{cases}$$
(7)

B. The evaporation cross section

By multiplying three quantities, the capture cross section, fusion probability, and survival probability, we can calculate the evaporation cross section for a superheavy element as [37]

$$\sigma_{\rm ER}^{xn}(E_{\rm cm}) = \sigma_{cap}(E_{\rm cm})P_{\rm CN}(E_{\rm cm})W_{\rm sur}^{xn}(E^*,l).$$
 (8)

This is an estimation because the total process of the compound nucleus formation and decay is separated here into three single reaction stages even if they are related to each other; in addition, they are treated and evaluated separately. In the fusion of heavy nuclei, it is the fission channels that considerably determine the dynamics of the total process; the fusion probability value can be significantly smaller than unity, whereas its exact calculation is difficult. Furthermore, currently, there is no agreement for the mechanism of the compound nucleus formation itself, and relatively different, occasionally opposite in their physics sense, models are used for the description.

In contrast to the excessively available data in the

study on the initial stage of the heavy ion fusion reactions and the processes of statistical decay of weakly excited compound nuclei, an accurate description of these reaction stages in the synthesis of superheavy elements is also difficult, which introduces an additional uncertainty in the cross section calculations of the evaporation residue formation. Here, the uncertainty is related with the convolution of the mechanisms of the first and last reaction stages, as well as with the fact that a number of quantities and nuclear specifications are not properly determined in this region.

The capture cross section can be evaluated using the sum of the cross section for each partial wave l as

$$\sigma_{\rm cap} = \frac{\pi}{k^2} \Sigma_{l=0}^{\infty} (2l+1) T(E,l), \tag{9}$$

where, $k = \sqrt{\frac{2\mu E}{\hbar^2}}$ and μ denote the reduced mass of the target and projectile systems. In addition, T(E,l) is the penetration probability of the *l*th partial wave, which can be calculated using the Hill-Wheeler formula [38]

$$T(E,l) = \left\{ 1 + \exp[2\pi(E_l - E)/\hbar\omega_l] \right\}^{-1}.$$
 (10)

Substituting the penetration probability in Eq. (9) by an integral, Wong obtained the following relationship for the capture cross section [39]

$$\sigma_{\rm cap} = \frac{R_0^2 \hbar \omega_0}{2E} \ln \left\{ 1 + \exp[2\pi (E - E_0)/\hbar \omega_0] \right\}, \qquad (11)$$

here E_0 is the barrier height. R_0 is the barrier radius and can be calculated as

$$\left(\frac{\mathrm{d}V(r)}{\mathrm{d}r}\right)_{r=R_0} = 0. \tag{12}$$

The curvature of the potential can be calculated as

$$\hbar\omega_0 = \hbar \sqrt{\frac{1}{\mu} \left(\frac{\mathrm{d}^2 V(r)}{\mathrm{d}r^2}\right)_{r=R_0}}.$$
(13)

Based on Ref. [40], we evaluated the fusion probability by the following relationship:

$$P_{\rm CN} = \frac{\exp(-c(x_{\rm e} - x_{\rm thr}))}{1 + \exp\left[\frac{E_{\rm B}^* - E^*}{\Delta}\right]},$$
(14)

where Δ is an adjustable parameter. $E_{\rm B}^*$ is the excitation energy of the compound nucleus when the center of mass beam energy is equal to the Coulomb and proximity energy. $x_{\rm e}$ can be evaluated as

$$x_{\rm e} = \left[\frac{(Z^2/A)}{(Z^2/A)_{\rm crit}}\right] [1 - \alpha + \alpha f(K)],\tag{15}$$

here $\alpha = 1/3$ and $(Z^2/A)_{crit}$ is given by

$$(Z^2/A)_{\text{crit}} = 50.883 \left[1 - 1.7826 \left(\frac{N-Z}{A} \right)^2 \right],$$
 (16)

$$f(K) = \frac{4}{K^2 + K + \frac{1}{K^2} + \frac{1}{K}},$$
(17)

$$K = (A_P / A_T)^{1/3}.$$
 (18)

Equation (14) was proposed by Zagrebaev to investigate the cold fusion reactions; however, this equation has also been adopted in various studies for the hot fusion reactions. For the best fit to the cold fusion reaction, it is inferred that the values of c and x_e are 136.5 and 0.79, re-

spectively. For the hot fusion reaction, the best fit for $x_e \le 0.8$ is c = 104 and $x_{thr} = 0.69$; while for $x_e > 0.8$, the values are c = 82 and $x_{thr} = 0.69$ [41].

We calculated the value of the survival probability under the evaporation of the x neutron as [42]

$$W_{\rm sur} = P_{xn}(E_{\rm CN}^*) \prod_{i=0}^{i_{\rm max}=x} \left(\frac{\Gamma_n}{\Gamma_n + \Gamma_f}\right)_{i,E^*},\tag{19}$$

where the index *i* is the number of emitted neutron, and Γ_n and Γ_f present the decay widths of the neutron evaporation and fission, respectively. Vandenbosch and Huizenga suggested that Γ_n/Γ_f can be calculated as [42]

$$\frac{\Gamma_n}{\Gamma_f} = \frac{4A^{2/3}a_f(E^* - B_n)}{K_0 a_n [2a_n^{1/2}(E^* - B_f)^{1/2} - 1]} \\ \times \exp[2a_n^{1/2}(E^* - B_n)^{1/2} \\ - 2a_f^{1/2}(E^* - B_f)^{1/2}],$$
(20)

here k = 9.8 MeV, and $a_n = A/10$ represents the level density parameter for the neutron evaporation channel. The level density parameter for fission is $a_f = 1.1a_n$. In addition, B_n and B_f represent the neutron separation energy and fission barrier, respectively. B_f is evaluated using the shell correction as

$$B_f(E^*) = B_f^{\text{LD}} + S \exp(-E^*/E_D),$$
 (21)

where B_f^{LD} is the liquid drop fission barrier that is equal to zero for heavy elements with Z > 109, and the shell correction, *S*, is obtained from Ref. [43]. In addition, the shell damping energy, E_D , is evaluated as

$$E_D = 5.48A^{1/3} / (1 + 1.3A^{-1/3}).$$
 (22)

In Eq. (19), $P_{xn}(E_{CN}^*)$ is calculated based on the method presented in Refs. [42, 44, 45] as

$$P_{xn}(E_{\rm CN}^*) = P[x] - P[x+1], \tag{23}$$

where P(x) is the probability that at least x neutrons are evaporated at a given E^* and can be calculated by the following relationship

$$P[x] = 1 - \exp(-\Delta_x/T) \left[1 + \sum_{i=1}^{2x-3} \frac{(\Delta_x/T)^i}{i!} \right], \quad (24)$$

with

$$\Delta_x = E_{\rm CN}^{\star} - \sum_{k=1}^x B_k,\tag{25}$$

here B_k is the evaporation neutron separation energy and

$$T = \sqrt{\frac{E_{\rm CN}^{\star}}{(2a_n)}}.$$

III. RESULTS

We theoretically investigated the formation of the superheavy element ²⁹⁶119 using the Coulomb and proximity potential models. We selected 29 reactions with different targets and projectiles, including convenient half lives. The evaporation residue cross section of the 3n and 4n channels is calculated. The entrance channel influences the production of the superheavy element via the Qvalue and entrance asymmetry. The mass and neutron asymmetries are defined as $\alpha_A = (A_T - A_P)/A$ and $\alpha_N = (N_T - N_P)/N$, respectively. The effect of the entrance channel is studied in comparison with the evaporation residue cross sections for the reactions triggering the formation of the same compound nucleus. The calculated results are presented in Figs. 1 to 12 and Table 1.

In Fig. 1, we display the variation of the potential energy with the distance between the projectile and the target for the reactions ${}^{49-47}\text{Ti} + {}^{247-249}\text{Bk}$, ${}^{60-57}\text{Fe} +$ $^{236-239}$ Np, $^{44-42}$ Ca + $^{252-254}$ Es, and 55,54,52 Mn + 241,242,244 Pu. In these reactions, with an increase in the entrance channel (mass or neutron) asymmetry, the fusion barrier height increases; however, the barrier radius decreases. The capture cross section as a function of energy is plotted in Fig. 2 for these reactions. The maximum capture cross section is related to the reaction ⁴⁹Ti+²⁴⁷Bk. The vertical arrows demonstrate the excitation energies of $E^* = E_0 + Q$ for different reactions. The neutron asymmetry parameters for the reactions, ${}^{49}\text{Ti}+{}^{247}\text{Bk}$, ${}^{48}\text{Ti}+{}^{248}\text{Bk}$, and ${}^{47}\text{Ti}+{}^{249}\text{Bk}$, are 0.668, 0.675, and 0.682, respectively. Because the reaction ${}^{49}\text{Ti}+{}^{247}\text{Bk}$ has the highest Q value at energies below the Coulomb barrier, the capture cross section is larger. In addition, the neutron asymmetry via the barrier hight, barrier radius, and curvature of the potential influence the capture cross section. In other words, owing to the effects of the Q value and Coulomb barriers, the capture cross sections are different. By increasing the energy, the capture cross section increases at low energies and then reaches the maximum value at higher energies. The same results can be concluded for other reactions.

The probability of the compound nucleus formation with the excitation energy is calculated in Fig. 3. At lower energies, with the increase in the Q value, the probability of the compound nucleus formation increases. At higher energies, the obtained results for the formation probability are approximately close together. It should be noted that in Fig. 3, the reaction ⁴⁴Ca+²⁵²Es exhibits the highest fusion probability.

In Fig. 4, we have demonstrated the 3n evaporation residue cross sections as a function of energy for the reac-



Fig. 1. Calculated dependence of the potential energy on the distance between the projectile and target.



Fig. 2. Calculated dependence of the capture cross section on the energy. The arrows denote the positions of the corresponding Coulomb barriers.



Fig. 3. Fusion probability as a function of energy.

tions $^{49-47}$ Ti + $^{247-249}$ Bk, $^{60-57}$ Fe + $^{236-239}$ Np, $^{44-42}$ Ca + $^{252-254}$ Es, and 55,54,52 Mn + 241,242,244 Pu. For each of these reactions, the maximum value of the evaporation residue occurs for the case where the Q value is the highest. In addition, owing to the increase in the asymmetry parameter, the maximum cross section of the 3*n* evaporation residue is created at larger energies. We can observe that the evaporation residue cross sections reach the maximum value at the energy of approximately $E^* = E_0 + Q$. The obtained results for the 4*n* evaporation residue are presented in Fig. 5. The same results can be inferred from Fig. 5, in comparison with Fig. 4.

In the aforementioned reactions, the difference between the mass asymmetry parameters is low. Consequently, to investigate the role of mass asymmetry, we compared the obtained results for the reactions ${}^{44}Ca+{}^{252}Es$ and 60 Fe+ 236 Np. The mass asymmetry parameters for these reactions are 0.702 and 0.594, respectively. The calculated results are presented in Fig. 6. For the reaction $\frac{44}{222}$ ⁴Ca+²⁵²Es, the fusion probability is higher than that for the reaction 60 Fe+ 236 Np. In other words, for the reaction with larger asymmetry parameters, the fusion probability is higher than that for the reaction with lower asymmetry parameters. The capture cross section for the reaction $\frac{44}{252}$ 44 Ca+ 252 Es is lower than that for the reaction 60 Fe+ 236 Np. The barrier height and Q value for the reaction 44 Ca+ 252 Es are 206.747 and 168.723 MeV, respectively; however, for the reaction 60 Fe+ 236 Np, the barrier height and Q value are 248.424 and 222.583 MeV, respectively. Consequently, at energies below the Coulomb barrier, the capture cross section for the reaction ${}^{60}\text{Fe}+{}^{236}\text{Np}$ is larger than that for the reaction ${}^{44}\text{Ca}+{}^{252}\text{Es}$. The maximum σ_{3n} values for the reactions ${}^{44}\text{Ca}+{}^{252}\text{Es}$ and ${}^{60}\text{Fe}+{}^{236}\text{Np}$ are 169.97 and 11.0 MeV, respectively, whereas for the maximum σ_{4n} values, we have 42.15 and 0.03 fb, respectively. It can be inferred that for the combinations with the large discrepancy of mass asymmetry, the effect of the mass asymmetry on the evaporation residue cross section is elucidated.

In Fig. 7, for clarity, we demonstrated the results for the reaction ${}^{44}\text{Ca}+{}^{248}\text{Cm}$. Here, we presented the fusion probability, capture cross section, and survival probability. In addition, the 3n and 4n evaporation residue cross sections as a function of excitation energy are calculated. At energies lower than $E_0 + Q$, the evaporation residue cross section decreases because the capture cross section and fusion probability are small. At energies higher than $E_0 + Q$, the evaporation residue cross section decreases because the survival probability decreases.

To better investigate the influence of the Q value, we demonstrated the variations in the maximum value of the evaporation residue cross section with the |Q| in Fig. 8. As observed from Fig. 8, with an increase in |Q|, the maximum value of the evaporation residue cross section for the 3n and 4n channels increases. In this study, the Q value.



Fig. 4. Plot of the 3*n* evaporation residue cross sections as a function of energy. The arrows indicate the positions of the corresponding Coulomb barriers.



Fig. 5. Same as Fig. 3 but for the 4*n* channel.



Fig. 6. Comparison of the results for the reactions ${\rm ^{44}Ca+^{252}Es}$ and ${\rm ^{60}Fe+^{236}Np}.$



Fig. 7. Plot of the survival probability, fusion probability, capture cross section, and the 3n and 4n evaporation residue cross section as a function of energy for the reaction ${}^{48}\text{Ca}{+}^{248}\text{Cm}$. Solid and dashed curves are results for the 3n and 4n channels, respectively. Experimental data are obtained from Ref. [56].

ues are obtained from Ref. [43]. To study the role of mass table on the evaporation residue cross section, we used the mass based on Ref. [46] for the reaction ⁴⁹Ti+²⁴⁷Bk. The obtained results are presented in Fig. 9. The |Q| values based on Refs. [43] and [46] are 187.62 and 182.19 MeV, respectively. It is clear that the Q value has an important effect on the evaporation residue cross section. The evaporation residue cross section based on the Q value of Ref. [43] is higher than that based on Ref. [46].

The obtained results for all reactions are presented in Table 1. Among all reactions, the ${}^{44}Ca+{}^{252}Es$ combination has the maximum value (169.97 fb) for the evaporation residue for the 3*n* channel. The asymmetry parameter and |Q| for this reaction are 0.662 and 192.012 MeV, respectively.

There are different processes in the heavy ion reactions that hinder the fusion process, such as deep inelastic, quasifission, and fast fission processes. Quasifission occurs when the dinuclear system starts to break down in-



Fig. 8. Plot of the maximum evaporation residue cross section versus Q-value of the reaction. Solid and open squares are results for the 3n and 4n channels, respectively. In addition, the solid and dashed lines represent liner fits on the theoretical results of 3n and 4n channels, respectively.

to two fragments without attaining the compound nucleus formation. In other words, the quasifission competes with the compound nucleus formation. The effective entrance channel fissility is defined as

$$\chi_{\text{eff}} = \frac{\frac{4Z_P Z_T}{(A_P^{1/3} + A_T^{1/3}) \times (A_P A_T)^{1/3}}}{50.883 \left[1 - 1.7826 \left(\frac{A - 2Z}{A}\right)^2\right]}.$$
 (26)

The mean fissility that indicates the appearance of the quasifission can be calculated as [47, 48]

$$\chi_m = 0.25\chi_{\rm CN} + 0.75\chi_{\rm eff},\tag{27}$$

where χ_{CN} is the compound nucleus fissility that indicates the stability of the compound nuclei against the fission as

$$\chi_{\rm CN} = \frac{Z^2/A}{50.883 \left[1 - 1.7826 \left(\frac{A - 2Z}{A}\right)^2\right]}.$$
 (28)



Fig. 9. Comparison of calculated results for different Q values. Solid and dashed curves are the obtained results using the Q value based on Ref. [43] and Ref. [46], respectively.

The reactions studied in this research have $\chi_{CN} = 1.009$. Consequently, the mean fissility depends solely on the effective entrance channel fissility. It is clear that the possibility of the quasifission depends on the asymmetry of the entrance channel. The critical asymmetry parameter is defined as

$$\alpha_{\rm as}^{\rm CR} = \begin{cases} 0 & \chi_{\rm CN} < 0.396, \\ 1.12 \sqrt{\frac{\chi_{\rm CN} - 0.396}{\chi_{\rm CN} - 0.156}} & \chi_{\rm CN} > 0.396. \end{cases}$$
(29)

If we have $\alpha < \alpha_{as}^{CR}$, the quasifission process can occur. Here, we have $\alpha_{as}^{CR} = 0.949$. Because α_{as}^{CR} is the same for the suggested reactions, we use another parameter (χ_m) to investigate the quasifission process. As mentioned in Ref. [48], the quasifission becomes primary for $\chi_m > 0.765$. We present the mean fissility for all reactions in Table 1. The lowest and highest values of χ_m are 0.7953 and 0.9229, related to the reactions $\frac{48}{22}$ Ti + $\frac{248}{97}$ Bk and $\frac{71}{31}$ Ga + $\frac{225}{88}$ Ra, respectively. Figure 10 presents the maximum values of the 3*n* and

Figure 10 presents the maximum values of the 3n and 4n evaporation residue cross sections as a function of mean fissility. It is important to note here that as mean fissility increases, the maximum value of evaporation residue cross section decreases. It can be inferred that



Fig. 10. Variations of the maximum evaporation residue cross section versus the mean fissility. Solid curves represent the second order fit for the theoretical results.

Table 1. Obtained results for 3*n*, 4*n* channels evaporation. Columns two, three, four, and five are $|Q| = |(M_P + M_T - M_{CN})c^2|$ [43], the mass asymmetry, neutron asymmetry, and mean fissility, respectively. The position of the maximum value of the 3*n* channel (E_{3n}^*) and the maximum value of the evaporation residue of the 3*n* channel (σ_{3n}^{\max}) are presented in the sixth and seventh columns. In addition, the results for the 4*n* channel (E_{4n}^*) are presented in columns eight and nine.

Reaction	Q /MeV	$\alpha_A = (A_T - A_P)/A$	$\alpha_N = (N_T - N_P)/N$	Χm	E_{3n}^*/MeV	$\sigma_{3n}^{\max}/{\rm fb}$	$E_{4n}^*/{ m MeV}$	$\sigma_{4n}^{\max}/{\rm fb}$
$^{38}_{18}\text{Ar} + ^{258}_{101}\text{Md} \rightarrow ^{296} 119$	147.576	0.7432	0.7740	0.8049	45.26	54.26	46.17	125.84
$^{36}_{18}\text{Ar} + ^{260}_{101}\text{Md} \rightarrow ^{296} 119$	138.229	0.7567	0.7966	0.8160	55.92	0.03	58.13	0.27
$^{39}_{19}\text{K} + ^{257}_{100}\text{Fm} \rightarrow ^{296} 119$	149.766	0.7364	0.7740	0.8243	52.02	0.07	55.32	0.54
$^{44}_{20}$ Ca + $^{252}_{99}$ Es \rightarrow^{296} 119	168.723	0.7027	0.7288	0.8230	39.36	169.97	41.82	42.15
$^{43}_{20}$ Ca $+^{253}_{99}$ Es \rightarrow^{296} 119	163.944	0.7094	0.7401	0.8277	44.63	2.92	45.59	5.77
$^{42}_{20}$ Ca $+^{254}_{99}$ Es \rightarrow^{296} 119	161.105	0.7162	0.7514	0.8326	48.01	0.24	48.57	0.93
$^{46}_{21}$ Sc + $^{250}_{98}$ Cf \rightarrow^{296} 119	175.138	0.6891	0.7175	0.8369	40.71	25.64	42.61	4.74
$^{45}_{21}$ Sc + $^{251}_{98}$ Cf \rightarrow 296 119	171.484	0.6959	0.7288	0.8414	44.86	2.5	45.78	1.44
$^{49}_{22}$ Ti $+^{247}_{97}$ Bk \rightarrow^{296} 119	187.62	0.6689	0.6949	0.8445	35.59	106.3	40.17	4.7
$^{48}_{22}$ Ti $+^{248}_{97}$ Bk \rightarrow^{296} 119	184.96	0.6756	0.7062	0.7953	38.49	10.5	41.35	2.2
$^{47}_{22}$ Ti $+^{249}_{97}$ Bk \rightarrow^{296} 119	179.635	0.6824	0.7175	0.8012	44.25	0.12	45.27	0.22
$^{51}_{23}$ V + $^{245}_{96}$ Cm \rightarrow^{296} 119	194.295	0.6554	0.6838	0.8566	34.82	27.64	39.93	0.53
$^{50}_{23}$ V + $^{246}_{96}$ Cm \rightarrow^{296} 119	193.332	0.6621	0.6949	0.8607	39.61	10.62	41.93	0.4
$^{54}_{24}\text{Cr} + ^{242}_{95}\text{Am} \rightarrow ^{296}$ 119	206.014	0.6351	0.6610	0.8642	32.63	70.34	39.06	0.21
$^{53}_{24}\text{Cr} + ^{243}_{95}\text{Am} \rightarrow ^{296}119$	202.659	0.6418	0.6723	0.8701	35.08	12.02	40.13	0.13
$^{55}_{25}$ Mn + $^{241}_{94}$ Pu \rightarrow^{296} 119	209.304	0.6283	0.6610	0.8791	35.31	0.98	40.09	0.04
$^{54}_{25}$ Mn + $^{242}_{94}$ Pu \rightarrow^{296} 119	205.388	0.6351	0.6723	0.8830	39.41	0.04	41.82	0.013
${}^{52}_{25}\text{Mn} + {}^{244}_{94}\text{Pu} \rightarrow {}^{296}119$	195.450	0.6486	0.6949	0.8912	50.37	1.02×10^{-5}	50.91	6×10^{-5}
${}^{60}_{26}\text{Fe} + {}^{236}_{93}\text{Np} \rightarrow {}^{296}119$	222.583	0.5945	0.6185	0.8789	30.55	11.0	38.00	0.03
${}^{59}_{26}$ Fe $+{}^{237}_{93}$ Np $\rightarrow {}^{296}$ 119	220.340	0.6013	0.6271	0.8823	31.46	3.4	38.9	0.02
${}^{58}_{26}$ Fe $+{}^{238}_{93}$ Np $\rightarrow {}^{296}$ 119	219.248	0.6081	0.6384	0.8859	32.55	1.8	39.2	0.015
$^{57}_{26}$ Fe + $^{239}_{93}$ Np \rightarrow^{296} 119	215.418	0.6148	0.6497	0.8896	34.95	0.28	40.3	0.007
$^{60}_{27}$ Co + $^{235}_{92}$ U \rightarrow^{296} 119	223.753	0.5912	0.6214	0.8975	33.91	0.27	39.61	0.0016
$^{59}_{27}$ Co $+^{237}_{92}$ U \rightarrow^{296} 119	221.387	0.6013	0.6384	0.8996	36.3	0.05	40.41	0.001
$^{64}_{28}$ Ni + $^{232}_{91}$ Pa \rightarrow 296 119	235.700	0.5675	0.5932	0.8991	30.83	0.49	38.93	7.3×10^{-4}
$^{63}_{28}$ Ni + $^{233}_{91}$ Pa \rightarrow 296 119	232.572	0.5743	0.6045	0.9023	31.98	0.02	38.31	5.3×10^{-4}
$^{65}_{29}$ Cu + $^{231}_{90}$ Th \rightarrow^{296} 119	237.995	0.5608	0.5932	0.9116	32.42	0.04	39.1	1.2×10^{-4}
$^{70}_{30}$ Zn + $^{226}_{89}$ Ac \rightarrow^{296} 119	249.804	0.5270	0.5480	0.9114	30.19	0.50	37.68	5.7×10^{-5}
$^{71}_{31}$ Ga + $^{225}_{88}$ Ra \rightarrow 296 119	252.695	0.5202	0.5480	0.9229	30.90	0.007	8.33	2×10^{-5}

with an increase in the mean fissility, the quasifission becomes dominant. We obtained the following relationships for evaporation residue cross sections versus the mean fissility as

$$\sigma_{3n}^{\max} = 10^{-310.23\chi_m^2 + 497.467\chi_m - 197.58},$$
(30)

and

$$\sigma_{4n}^{\max} = 10^{-274.834\chi_m^2 + 435.622\chi_m - 171.211}.$$
 (31)

We checked Eqs. (30) and (31) for the reactions ${}^{50}\text{Ti}+{}^{249}\text{Bk}$, ${}^{48}\text{Ca}+{}^{252}\text{Es}$, ${}^{48}\text{Ca}+{}^{254}\text{Es}$, ${}^{47}\text{Ti}+{}^{249}\text{Cf}$, and ${}^{50}\text{Ti}+{}^{249}\text{Cf}$. The obtained results for 3*n* channel are 25.92, 70.20, 70.07, 13.36, and 24.77 fb, while the results for the 4*n* channel are 6.46, 23.68, 23.47, 3.16, and 6.14 fb. Eqs. (30) and (31) are important because they determine the relationship between the maximum evaporation cross section and the quasifission process (via mean fissility).

It should be noted that the synthesis of superheavy elements is influenced by dynamical effects such as friction, inertia mass, and the dissipation of the kinetic energy. These effects have been investigated in several studies [49-52]. The formation of a superheavy nucleus can be dynamically evaluated using multidimensional Langevin equations. Nuclear dissipation delays fission, which enhances prescission particle multiplicities and evaporation residue cross sections relative to the standard statistical model predictions. In Ref. [52], it is inferred that the evaporation residue cross section is a very sensitive probe for nuclear friction. The decay of superheavy elements can be evaluated using statistical models,



Fig. 11. Comparison between the theoretical results of the 3n and 4n channels with the experimental data (solid and open squares) [11, 54, 55] for the reactions ${}^{26}Mg+{}^{248}Cm, {}^{48}Ca+{}^{238}U$, and ${}^{48}Ca+{}^{240}Pu$.



Fig. 12. Same as Fig. 11 but for the reactions ${}^{48}\text{Ca}+{}^{237}\text{Np}$, ${}^{48}\text{Ca}+{}^{242,244}\text{Pu}$, and ${}^{48}\text{Ca}+{}^{249}\text{Bk}$. Experimental data are obtained from [10, 55-57], for the reactions ${}^{26}\text{Mg}+{}^{248}\text{Cm}$, ${}^{48}\text{Ca}+{}^{238}\text{U}$, and ${}^{48}\text{Ca}+{}^{240}\text{Pu}$.

where the Kramers factor plays a decisive role [30, 53]. The Kramers factor depends on the nuclear friction coefficient, and it is inferred as a constraint in the phase space around the saddle point owing to friction.

We checked the reliability of the model by comparing the evaporation residue cross section and the experimental data for the reactions ${}^{26}Mg+{}^{248}Cm$, ${}^{48}Ca+{}^{238}U$, ${}^{48}Ca+{}^{240}Pu$, ${}^{48}Ca+{}^{237}Np$, ${}^{48}Ca+{}^{242,244}Pu$, and ${}^{48}Ca+{}^{249}Bk$. The results of the 3*n* and 4*n* channels are presented in Figs. 11 and 12. The solid and open squares represent the experimental data for the 3*n* and 4*n* channels, respectively. The agreement between theoretical results and experimental data is satisfactory.

IV. SUMMARY AND CONCLUSION

We theoretically studied the effect of entrance channel on the production of supherheavy ²⁹⁶119 via the calculation of the evaporation residue cross sections for the 3n and 4n channels. The synthesis of superheavy elements significantly depends on the entrance channel effects. The asymmetry parameter and the Q value play important roles in the formation of superheavy elements. With an increase in the asymmetry parameter, the evaporation residue cross section of the 3n and 4n channels decreases. In addition, with an increase in |Q|, the evaporation residue cross section increases. There are linear relationships between $\log_{10}\sigma_{3n}^{max}$ and $\log_{10}\sigma_{4n}^{max}$ with the |Q|

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value, for the reactions ${}^{49-47}\text{Ti} + {}^{247-249}\text{Bk}$, ${}^{60-57}\text{Fe} + {}^{236-239}\text{Np}$, ${}^{44-42}\text{Ca} + {}^{252-254}\text{Es}$, and ${}^{55,54,52}\text{Mn} + {}^{241,242,244}\text{Pu}$. The major distinction is owing to the different capture and fusion probabilities.

The sensitivity of the evaporation residue cross section to the mass models was investigated. It was observed that a change in the Q value can trigger significant changes in the evaporation residue cross section. Hence, such changes must be considered while calculating the evaporation residue cross section.

We investigated the role of mass asymmetry using the comparison of results for the reactions ⁴⁴Ca+²⁵²Es and ⁶⁰Fe+²³⁶Np. It was inferred that the reactions with larger mass asymmetries are more desirable for fusion than those with smaller mass asymmetries.

It should be noted that in heavy ion reactions, the role of quasifission is important and should be considered. This inference is expressed by the mean fissility parameter. By increasing the mean fissility parameter, the fusion probability decreases. In other words, for the reactions with larger values of the mean fissility parameter, the probability of the quasifission process increases. The quasifission probability is significant as it facilitates the probability analysis of the compound nucleus formation in the synthesis of superheavy nuclei. In general, to study the formation of superheavy elements, it is crucial to consider the three quantities as the asymmetry parameters, Q value and mean fissility.

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